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Bistability of the climate around the habitable zone: a thermodynamic investigation

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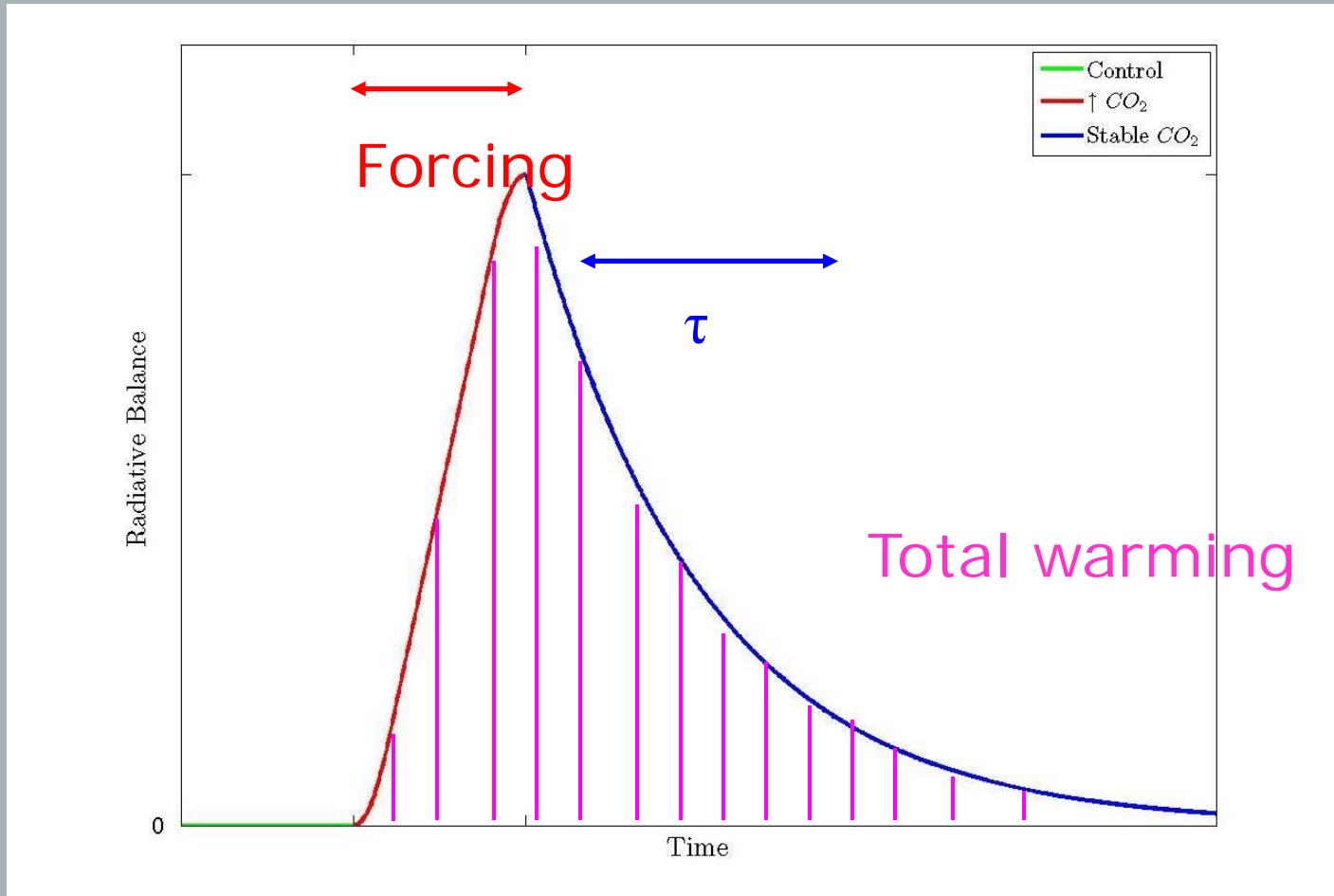
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R. Boschi, E. Kirk, N. Iro, S. Pascale, F. Ragone

Lorenz Center, 11/02/2014

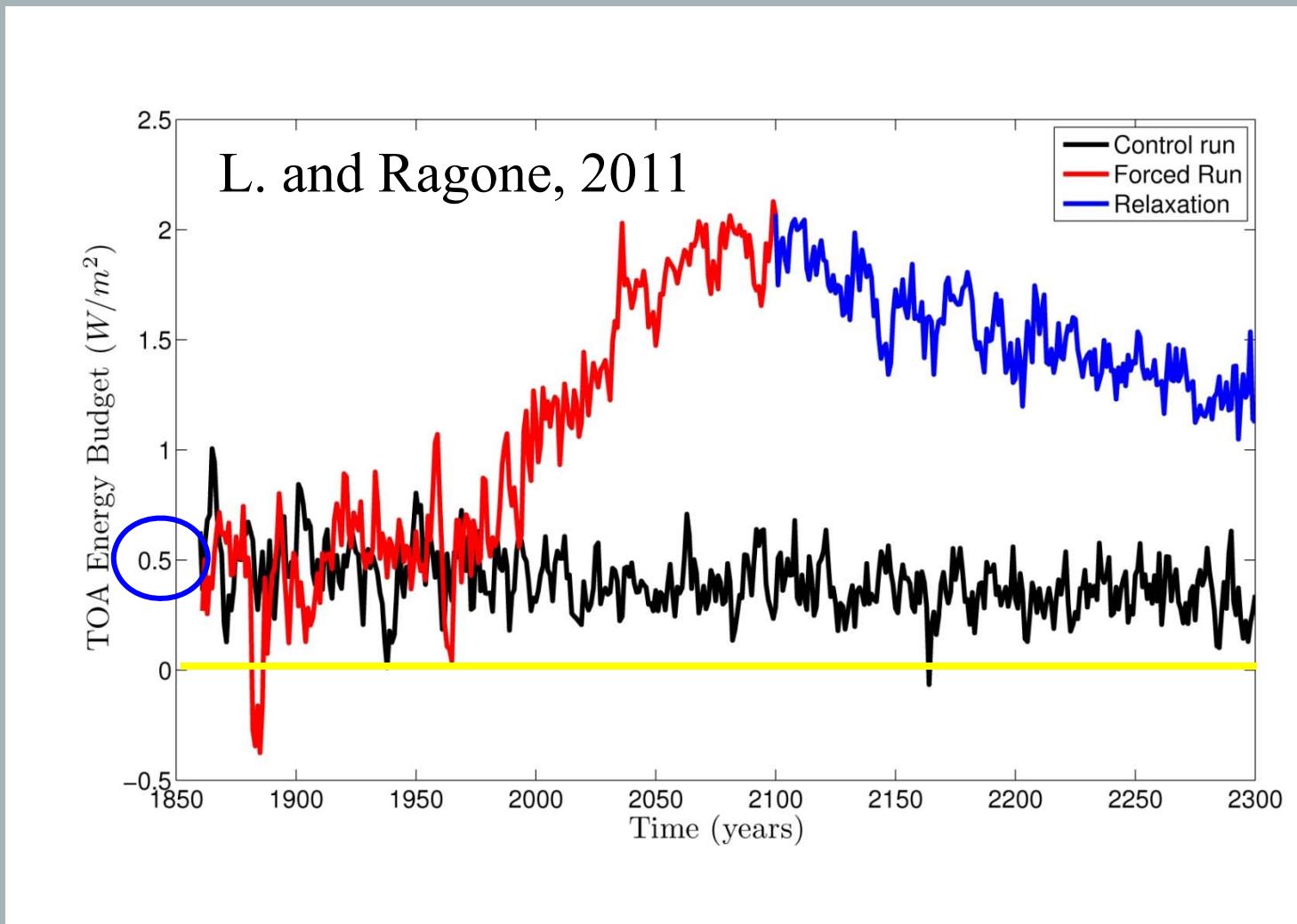
Energy & Forcing – Perfect Model



- ▲ NESS → Transient → NESS
- ▲ Applies to the whole climate and to all climatic subdomains
 - ▲ for atmosphere τ is small, always quasi-equilibrated ²

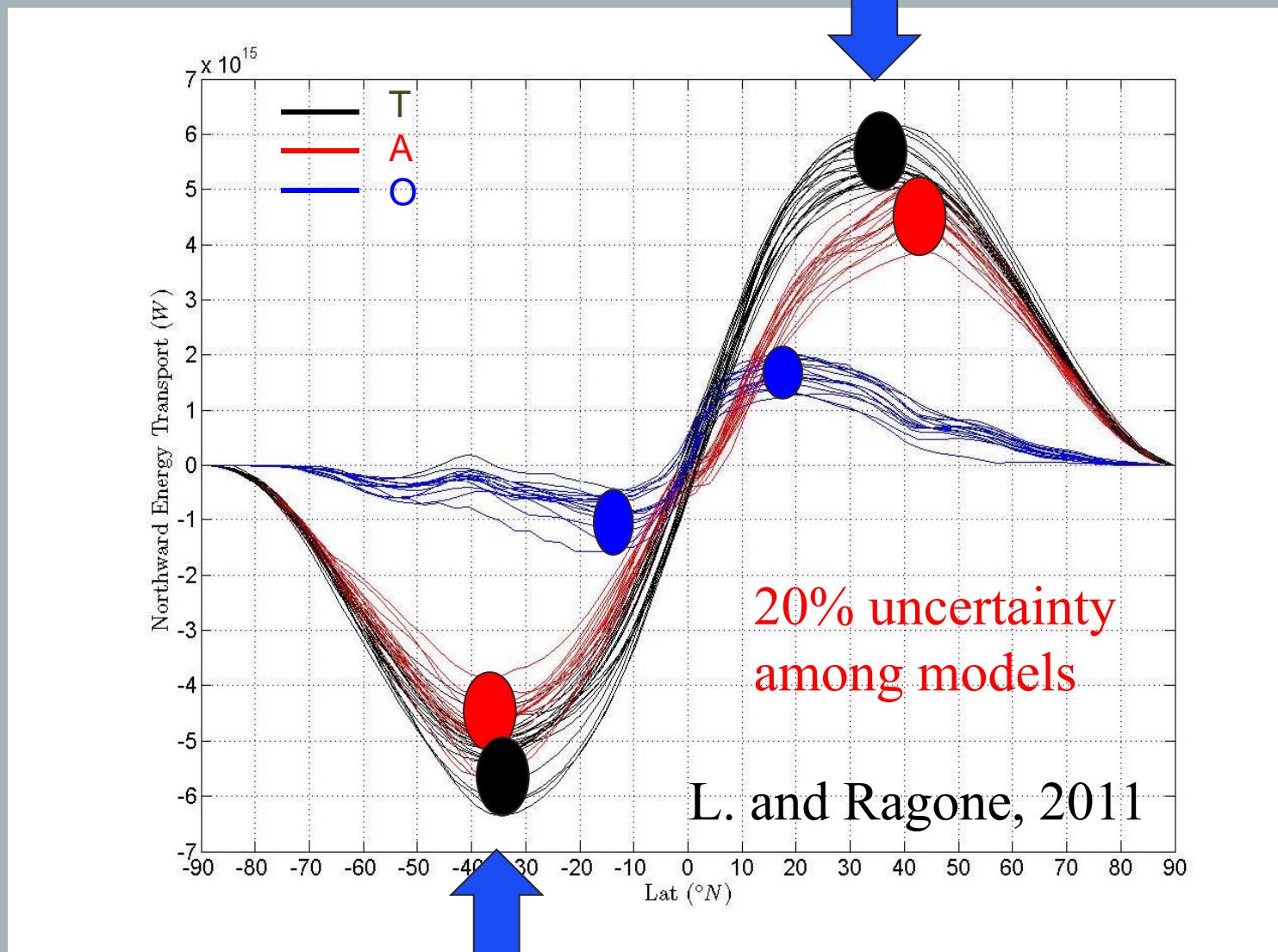


Energy and GW – Actual GCMs



- Not only bias: bias control \neq bias final state
Bias depends on climate state! \rightarrow Dissipation & Water 3

Steady State – Meridional Transports



Energies

▲ Kinetic energy budget

$$\dot{K}(\Omega) = - \int_{\Omega} dV \varepsilon^2 + C(P, K) = -D + W$$

WORK

$$W = C(P, K)$$

▲ Moist Static Potential Energy budget

$$\dot{P}(\Omega) = \int_{\Omega} dV \rho \dot{Q} - W \quad \dot{Q} = 1/\rho (\varepsilon^2 - \vec{\nabla} \cdot \vec{H})$$

▲ Total Energy Budget

$$\dot{E}(\Omega) = \int_{\Omega} dV (-\vec{\nabla} \cdot \vec{H}) = - \int_{\partial\Omega} dS \hat{n} \cdot \vec{H}$$

FLUXES

DISSIPATION

5

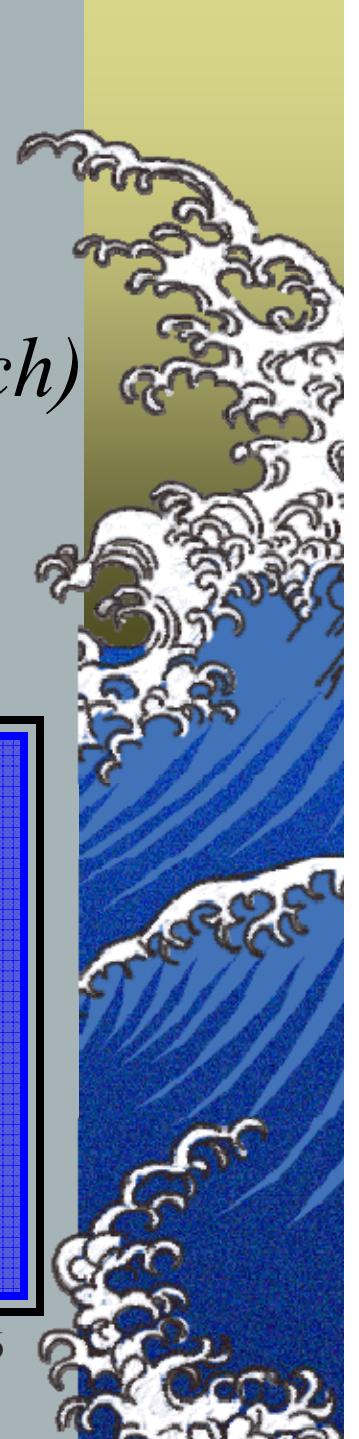
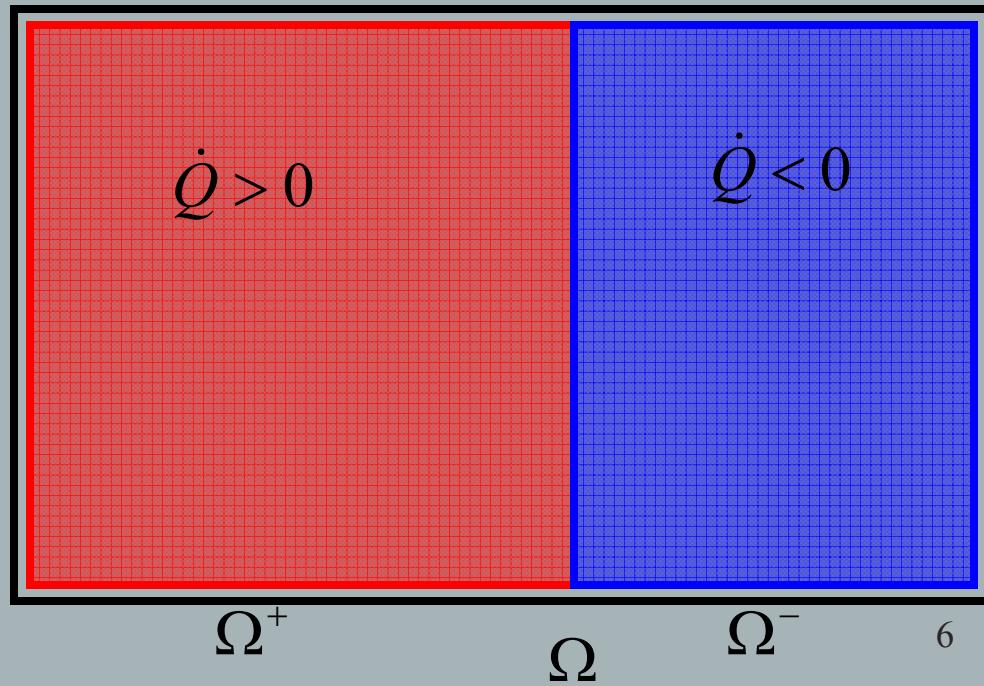


Johnson's idea (2000)

- ▶ *Partitioning the Domain (Eulerian approach)*

$$\dot{P}(\Omega) + W = \int_{\Omega^+} dV \rho \dot{Q}^+ + \int_{\Omega^-} dV \rho \dot{Q}^- = \dot{\Phi}^+ + \dot{\Phi}^-$$

- ▶ *Better than it seems!*



Long-Term averages

▲ *Stationarity:*

$$\overline{\dot{E}(\Omega)} = \overline{\dot{P}(\Omega)} = \overline{\dot{K}(\Omega)} = 0$$

▲ *Work = Dissipation*

$$-\overline{\dot{K}(\Omega)} + \overline{W} = \overline{W} = \overline{D} > 0$$

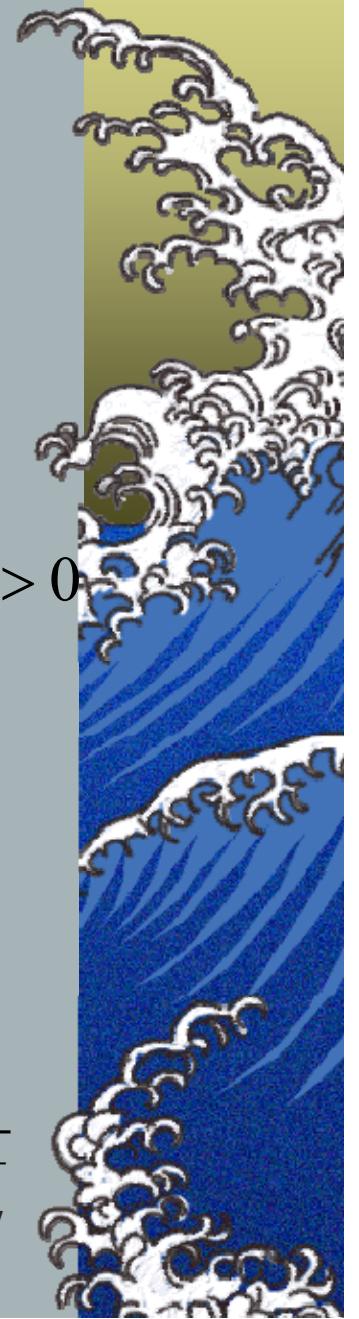
▲ *Work = Input-Output*

$$\overline{\dot{P}(\Omega)} + \overline{W} = \overline{W} = \overline{\dot{\Phi}^+} + \overline{\dot{\Phi}^-} > 0$$

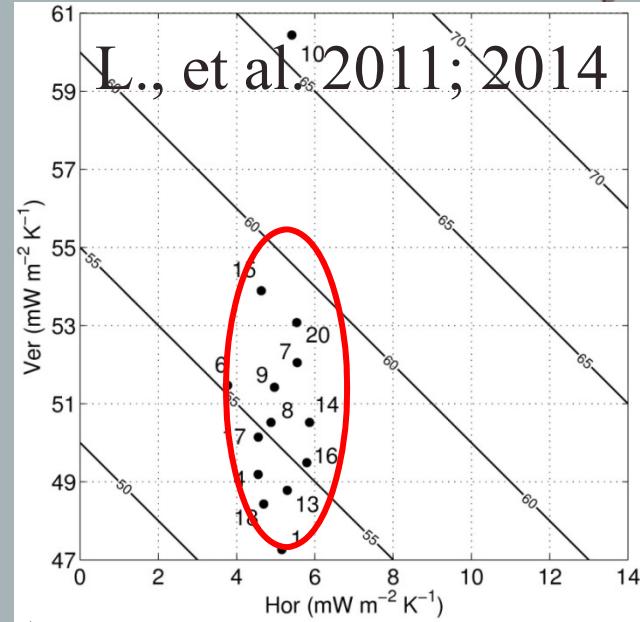
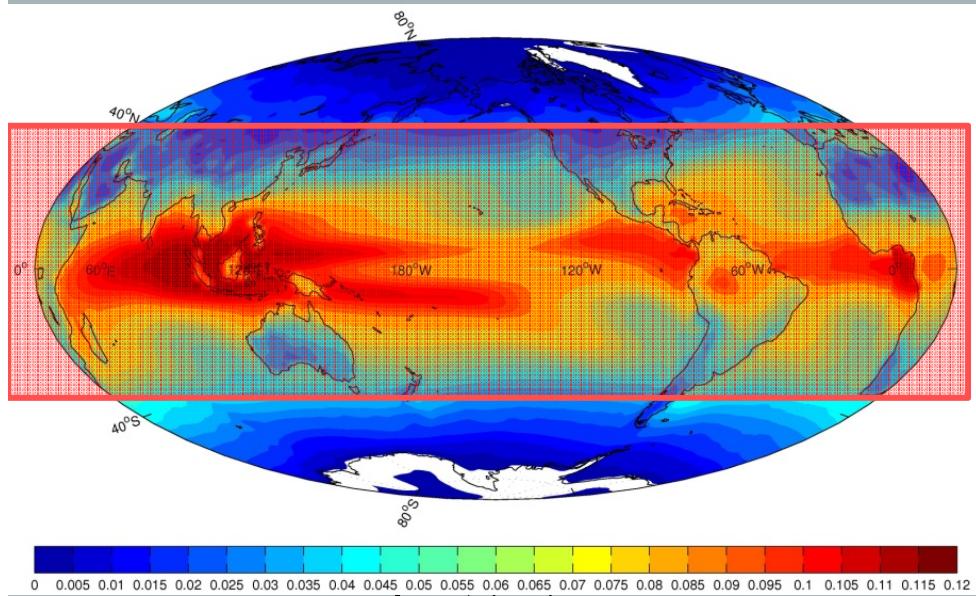
▲ *A different view on Lorenz Energy cycle*

$$\underbrace{\overline{\dot{\Phi}^+} + \overline{\dot{\Phi}^-}}_{\substack{\text{differential heating} \\ G(A)}} = \underbrace{\overline{W}}_{\substack{\text{conversion} \\ C(A,K)}} = \underbrace{\overline{D}}_{\substack{\text{dissipation} \\ D(K)}} > 0$$

$$\overline{W} = \frac{\overline{\dot{\Phi}^+} + \overline{\dot{\Phi}^-}}{\overline{\dot{\Phi}^+}} \overline{\dot{\Phi}^+} = \frac{\Theta^+ - \Theta^-}{\Theta^+} \overline{\dot{\Phi}^+} = \underbrace{\eta}_{\substack{\text{efficiency}}} \overline{\dot{\Phi}^+}$$

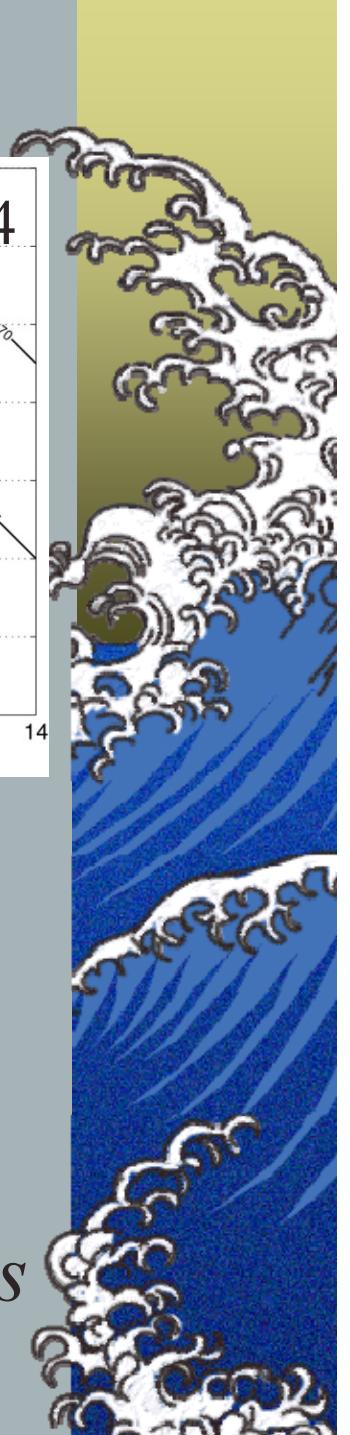


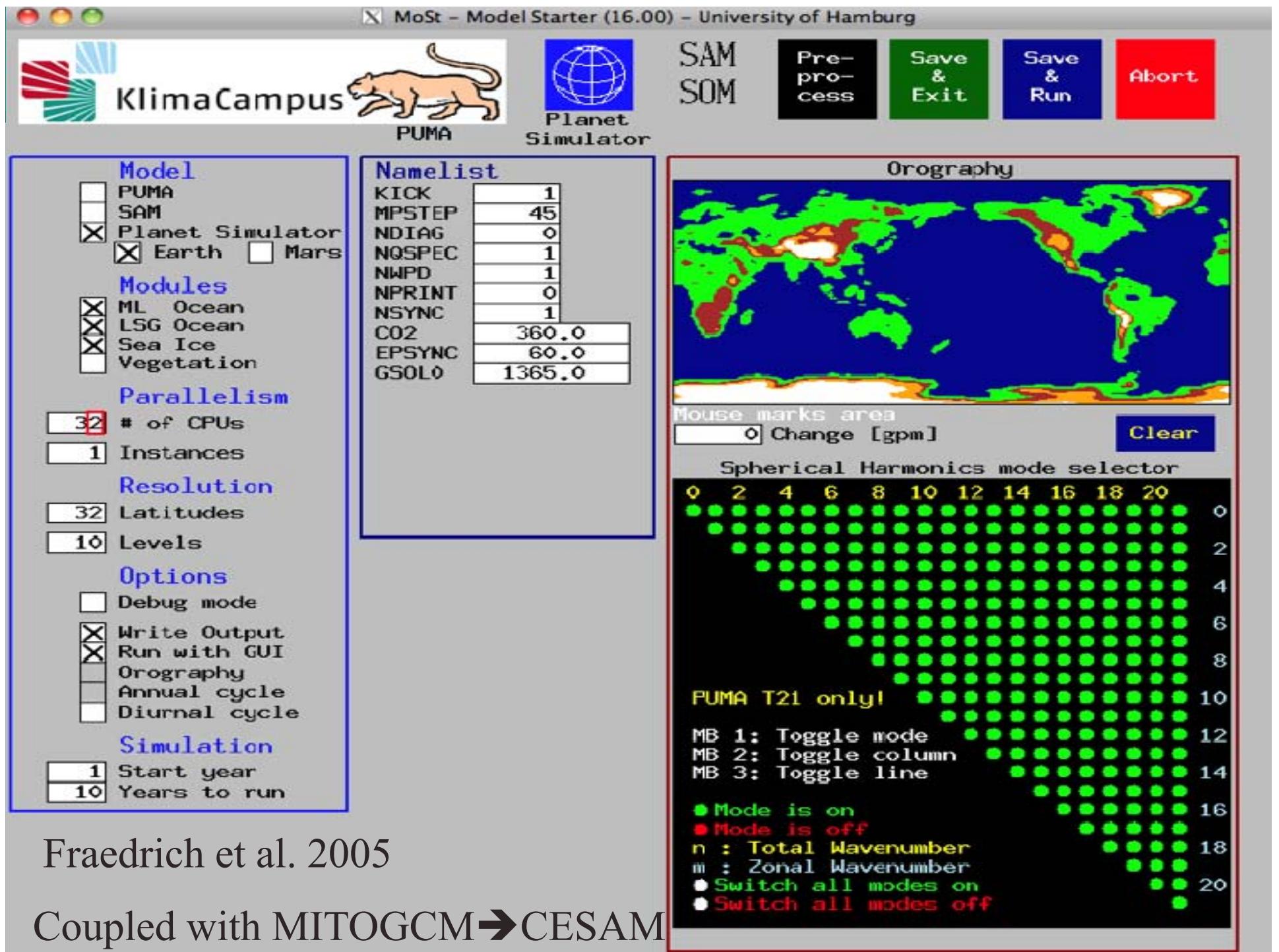
Results on IPCC GCMs



$$\frac{\dot{S}_{in}^{vert}(\Omega)}{\dot{S}_{mat}(\Omega)} \approx \underbrace{\int_S d\sigma R_{net}^{surf} \left(\frac{1}{T_E} - \frac{1}{T_S} \right)}_{\dot{S}_{mat}^{vert}(\Omega)} + \underbrace{\int_S d\sigma R_{net}^{TOA} \frac{1}{T_E}}_{\dot{S}_{mat}^{hor}(\Omega)}$$

- ▲ *Hor vs Vert EP in IPCC models*
- ▲ *Collection of (weak. coup.) vertical columns*
- ▲ *Warmer climate: Hor↓ Vert↑*

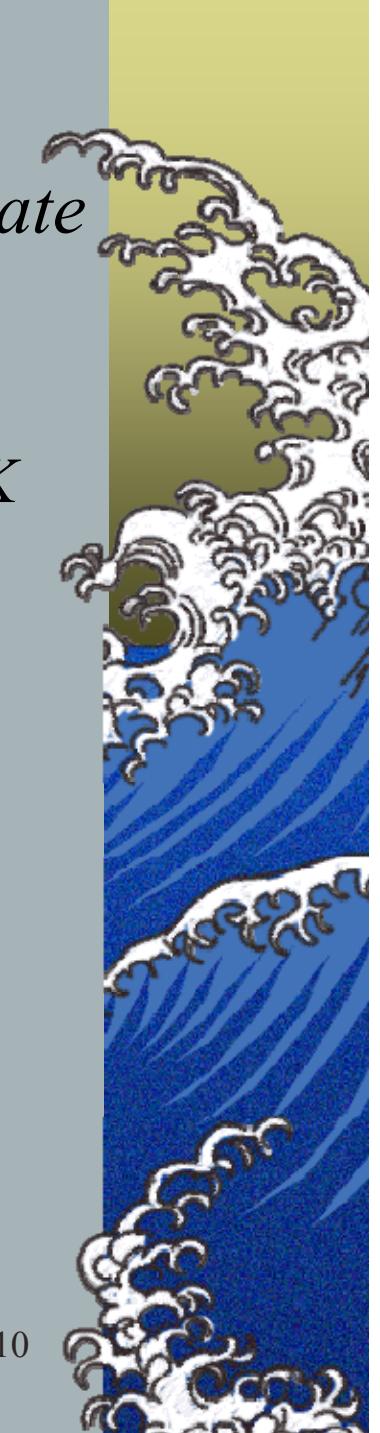
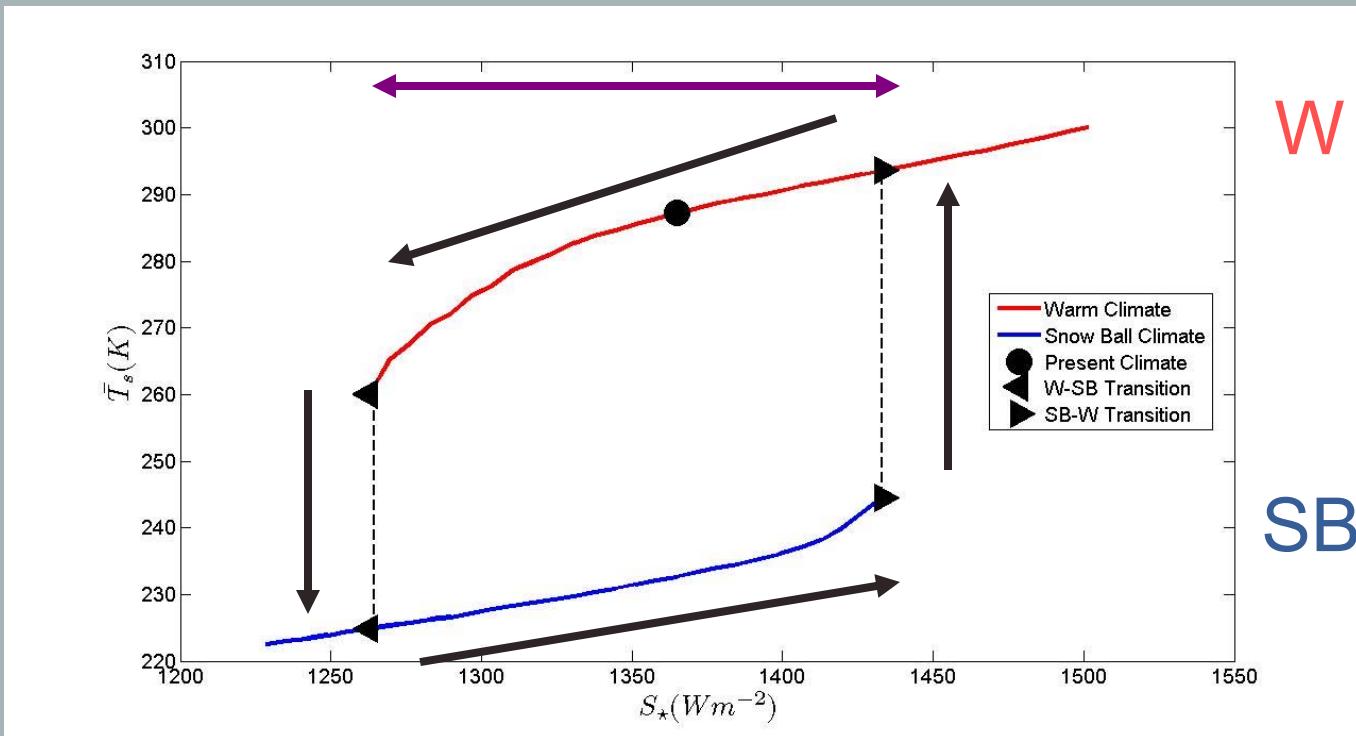




Snowball Hysteresis

- Swing of S^* by $\pm 10\%$ starting from present climate
 - → hysteresis experiment!
- Global average surface temperature T_s
 - Wide ($\sim 10\%$) range of S^* bistable regime - $\Delta T_s \sim 50 K$
 - $d T_s / d S_* > 0$ everywhere, almost linear

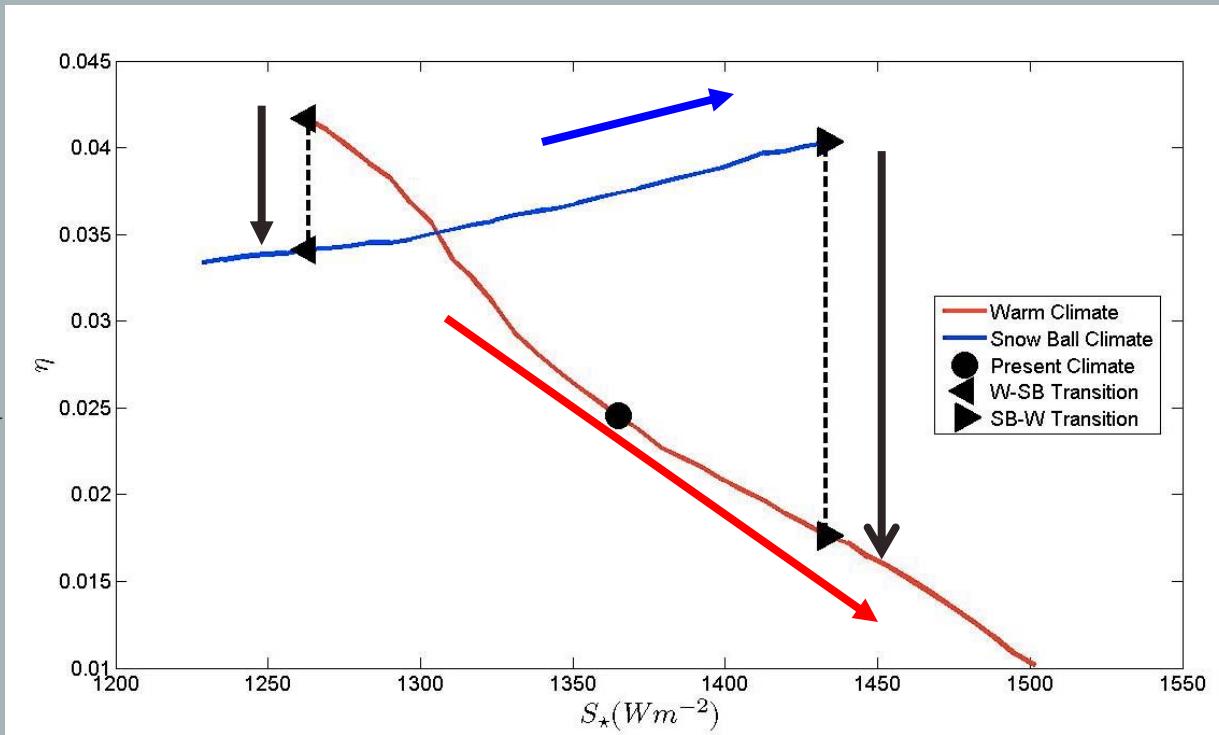
L., Lunkeit, Fraedrich, 2010



Thermodynamic Efficiency

- $d\eta/dS_* > 0$ in SB regime
 - Effect of decreased static stability
- $d\eta/dS_* < 0$ in W regime
 - System thermalized by efficient LH fluxes
- η decreases at transitions → System more stable

$$\eta = 0.04 \\ \leftrightarrow \\ \Delta\theta = 10K$$



Let's alter also $[\text{CO}_2]$

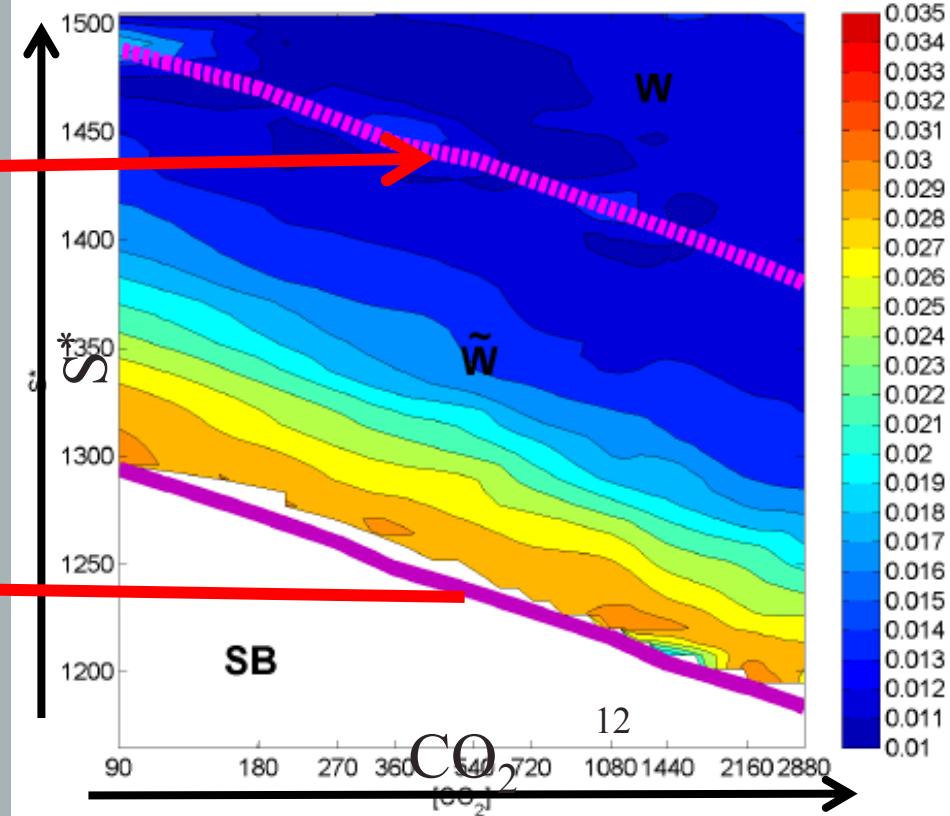
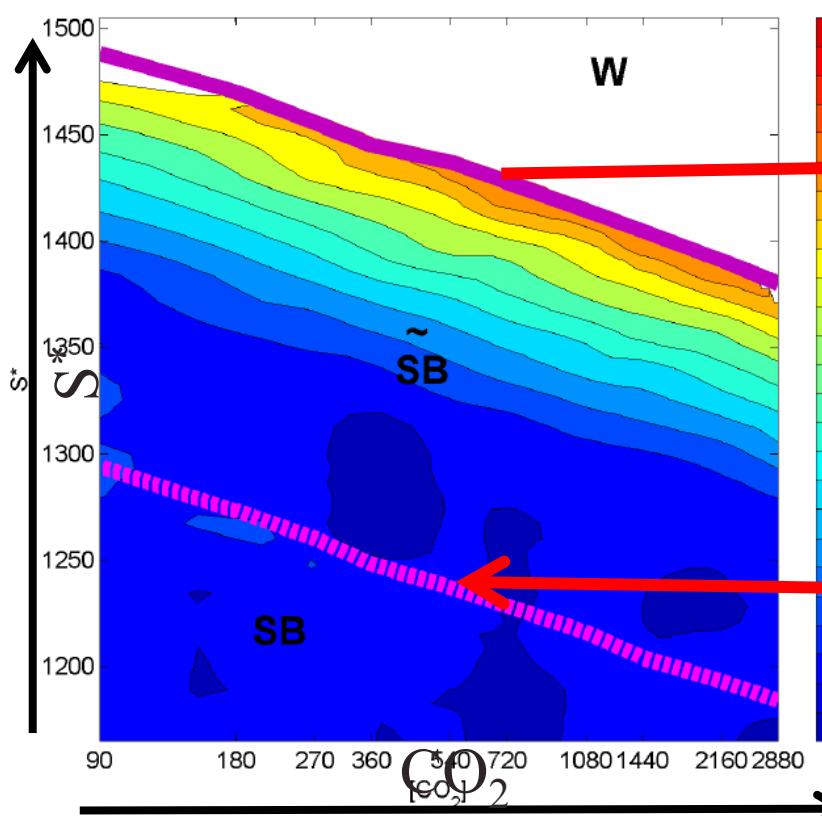
▲ *Parametric Analysis of Climate Change*

▲ *Structural Properties of the system (Boschi, et al. 2013)*

Lower Manifold

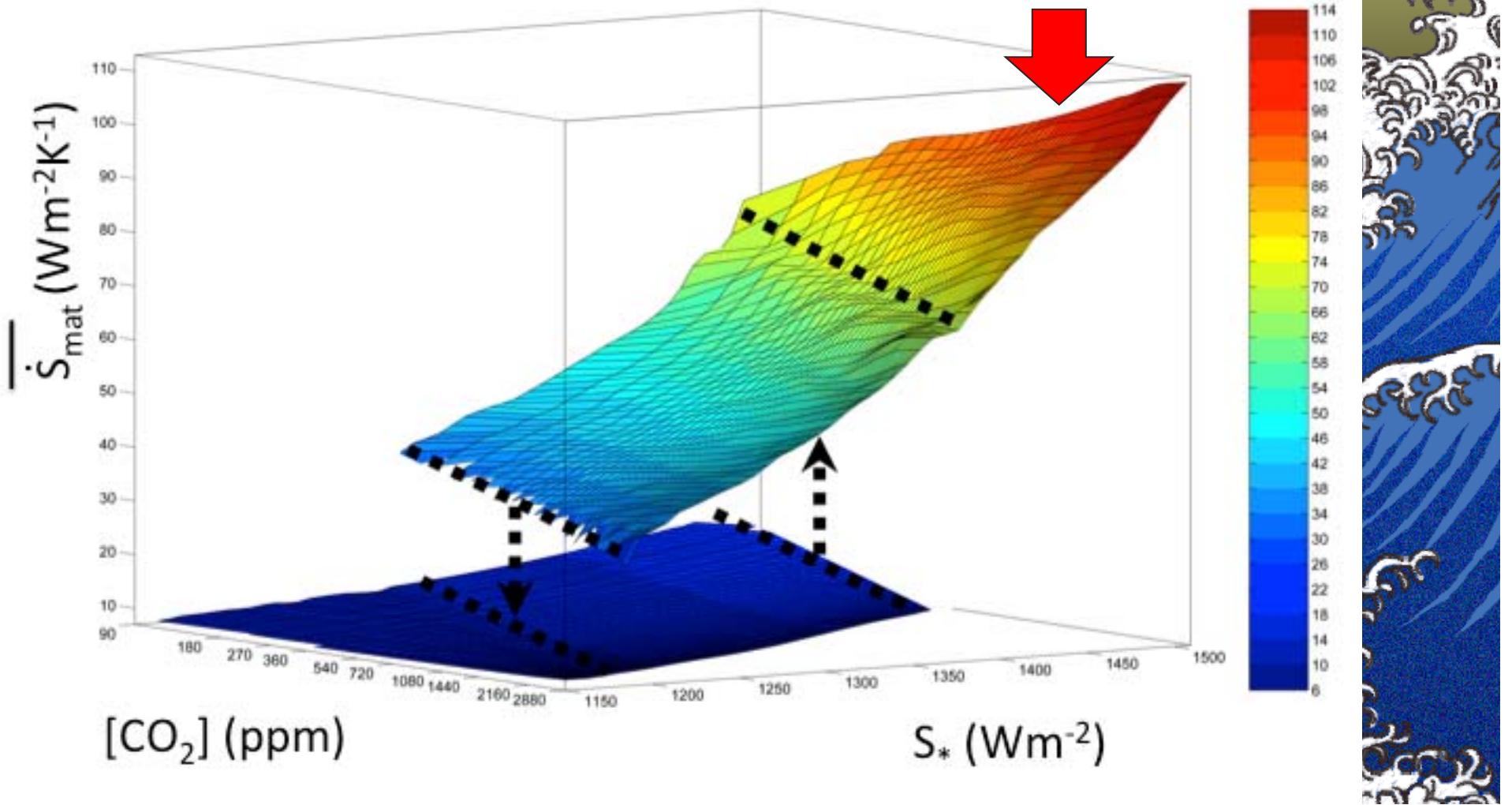
η

Upper Manifold



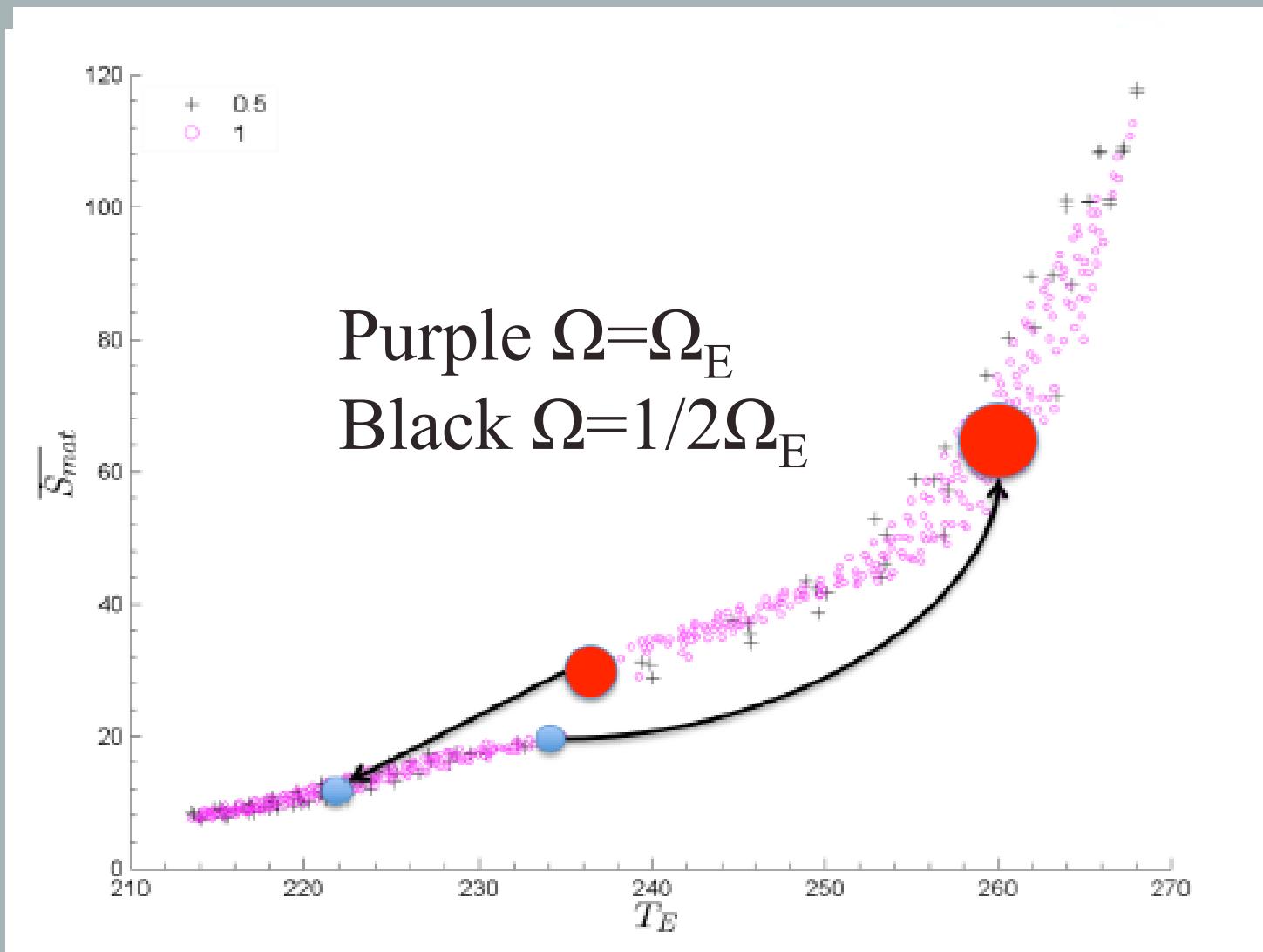
A 3D picture - EP

$$\overline{\dot{S}_{mat}(\Omega)} = \int_{\Omega} dV \vec{H} \cdot \vec{\nabla} \left(\frac{1}{T} \right) + \int_{\Omega} dV \frac{\varepsilon^2}{T} \quad Be = \overline{\dot{S}_{mat}(\Omega)} / \overline{\dot{S}_{min}(\Omega)}$$
$$= \overline{\dot{S}_{mat}(\Omega)} \langle \Theta \rangle / \bar{W}$$



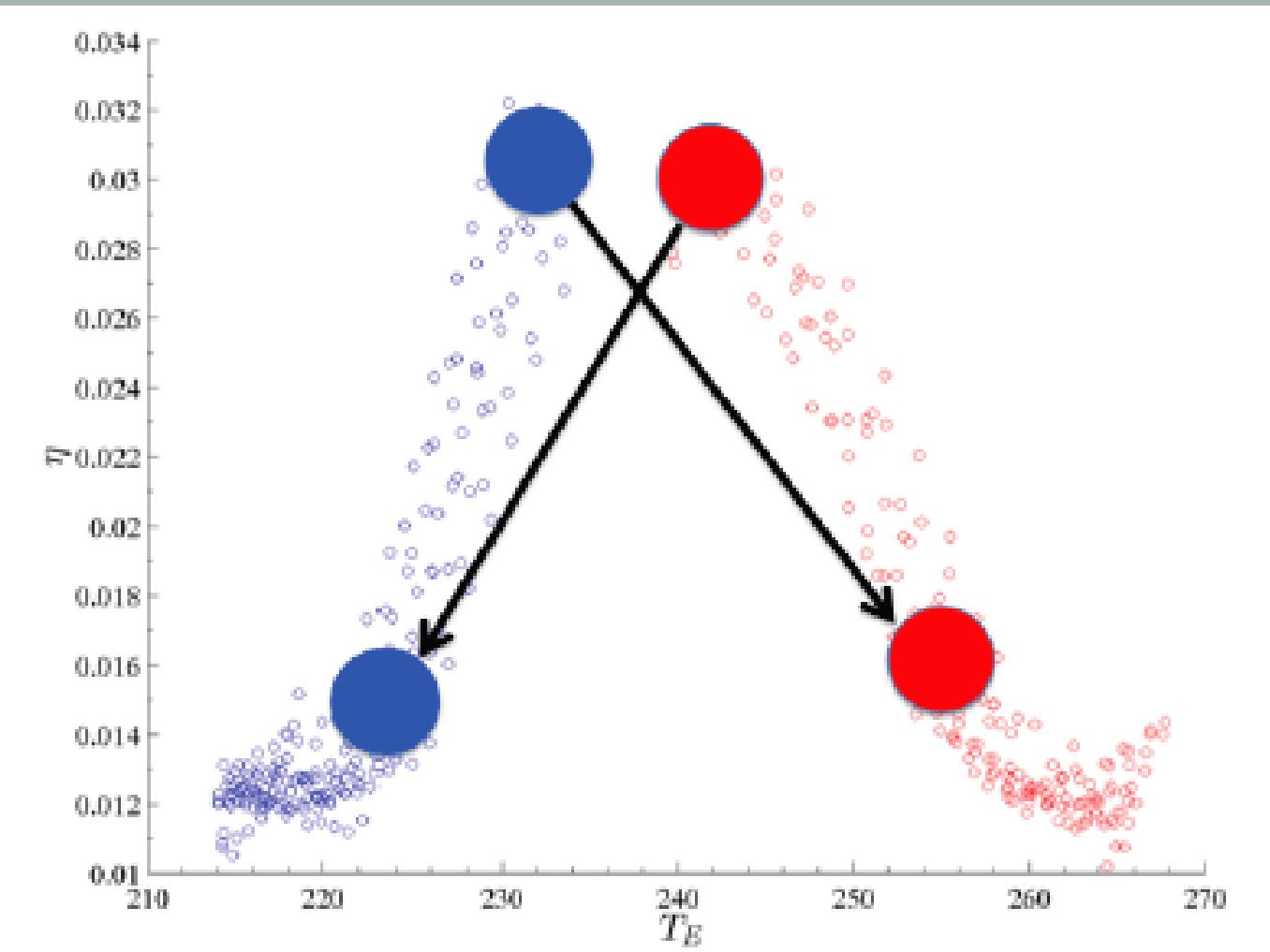
Parametrizations

▲ EP vs Emission Temperature



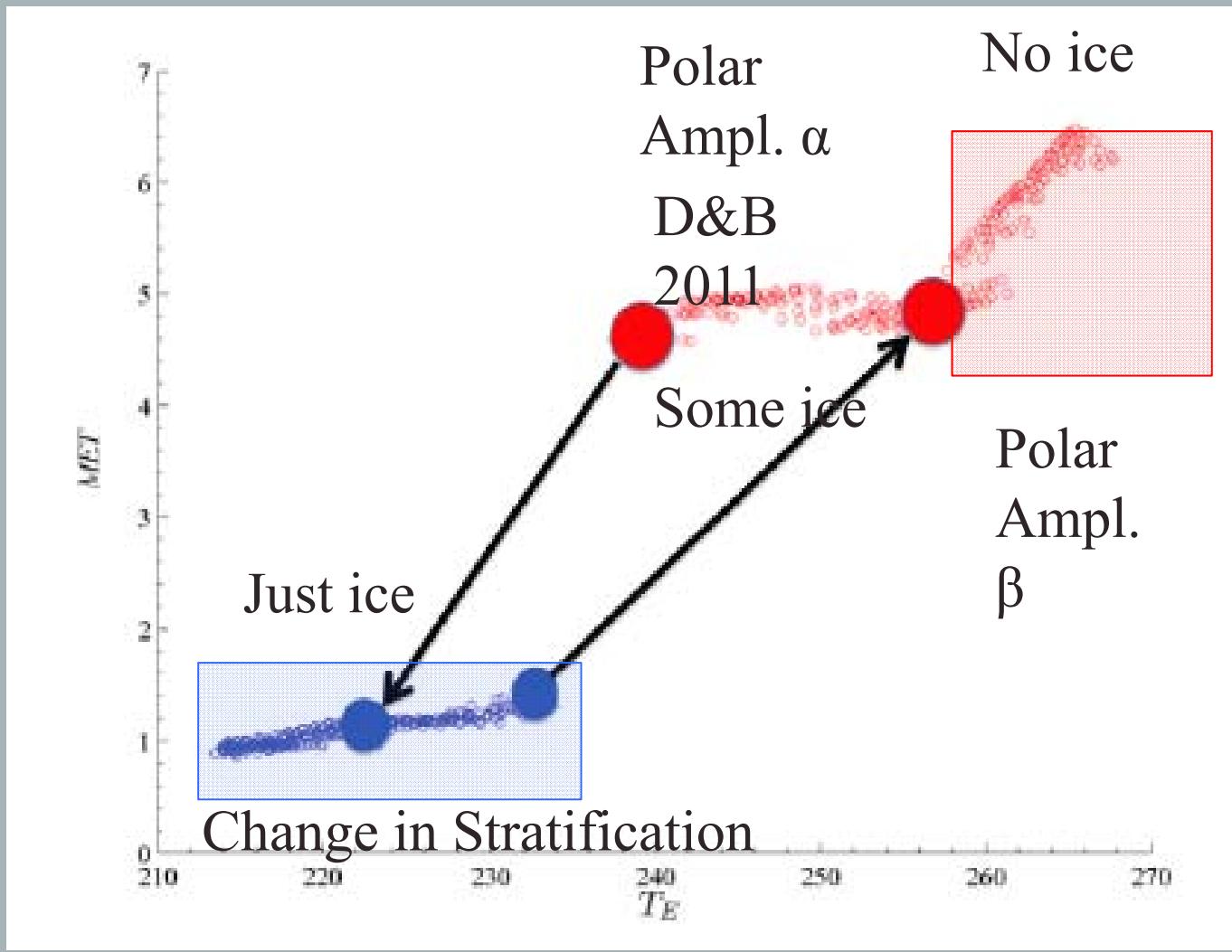
Parametrizations

▲ *Efficiency vs Emission Temperature*

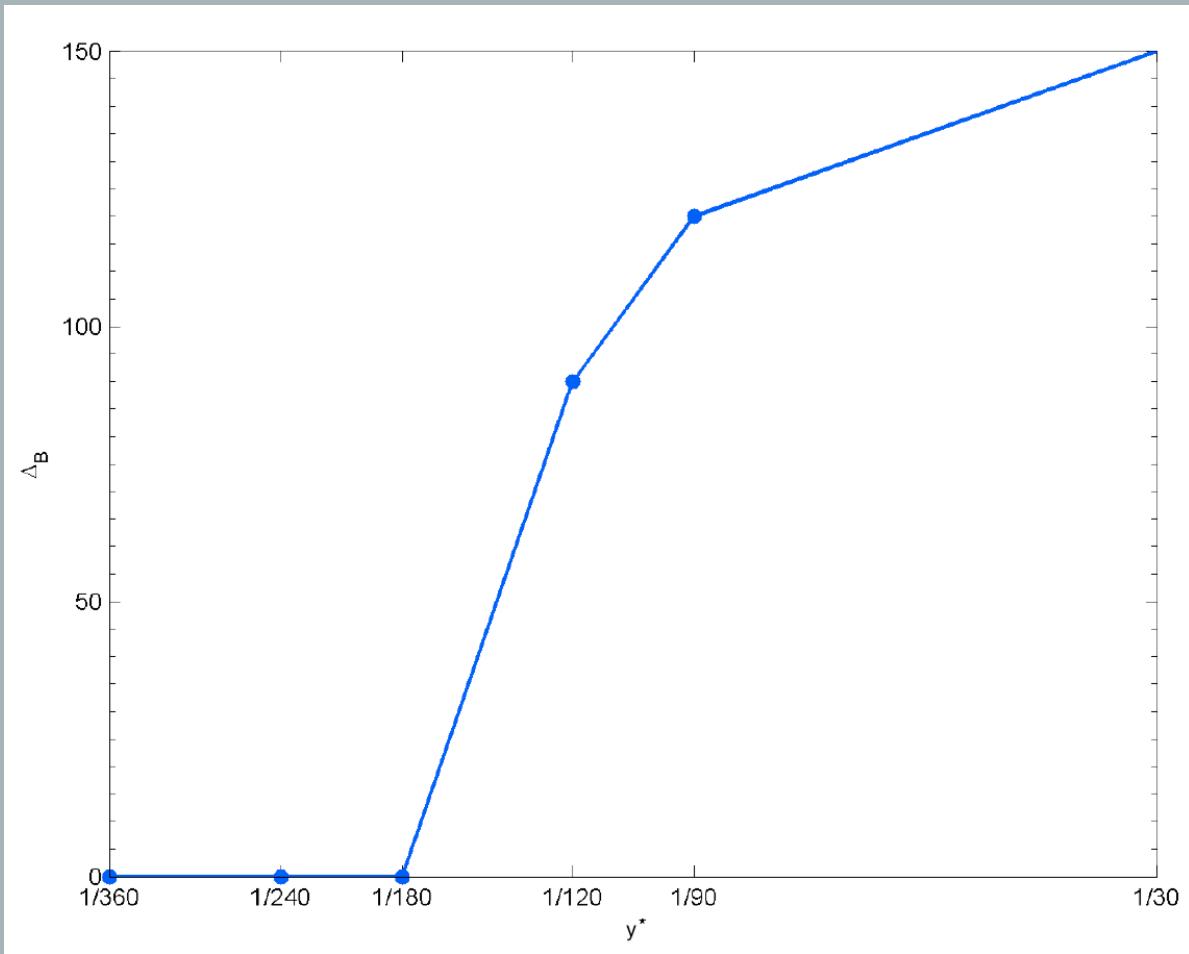


Parametrizations

▲ Heat Transport vs Emission Temperature

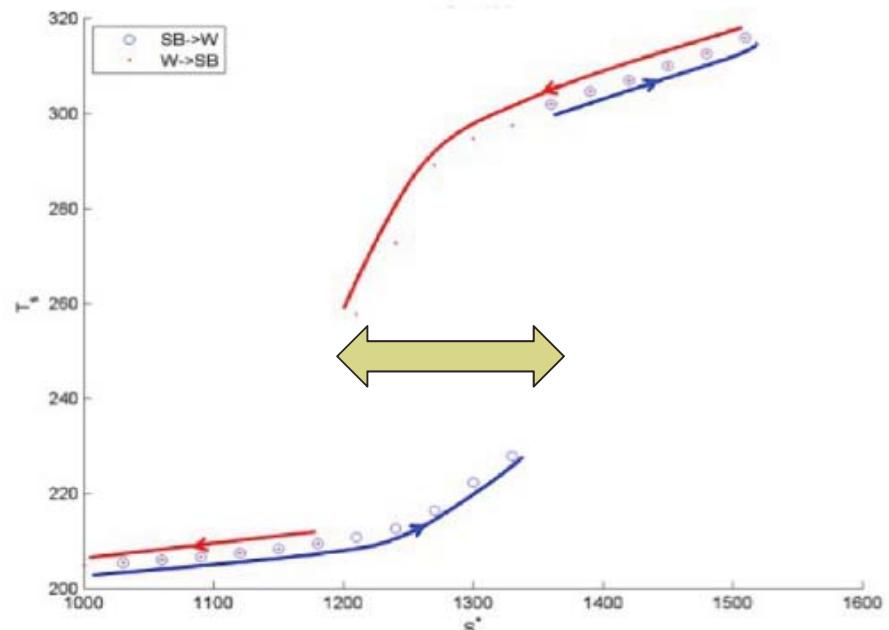


Shorter year - Phase Transition

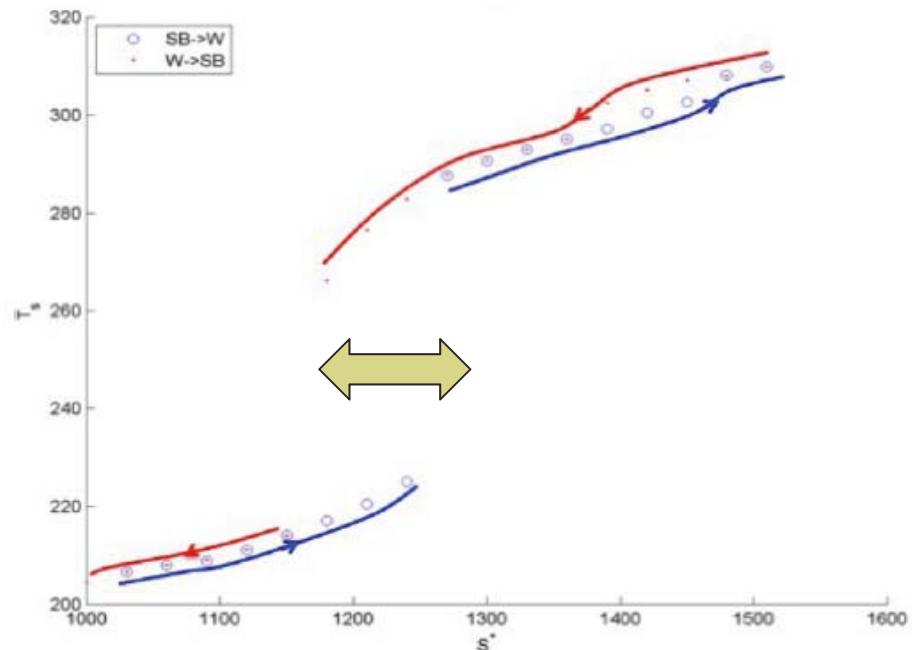


- Width bistability vs length year (L. et al. 2013)
 - Fast orbiting planets cannot be in Snowball Earth
 - Habitability

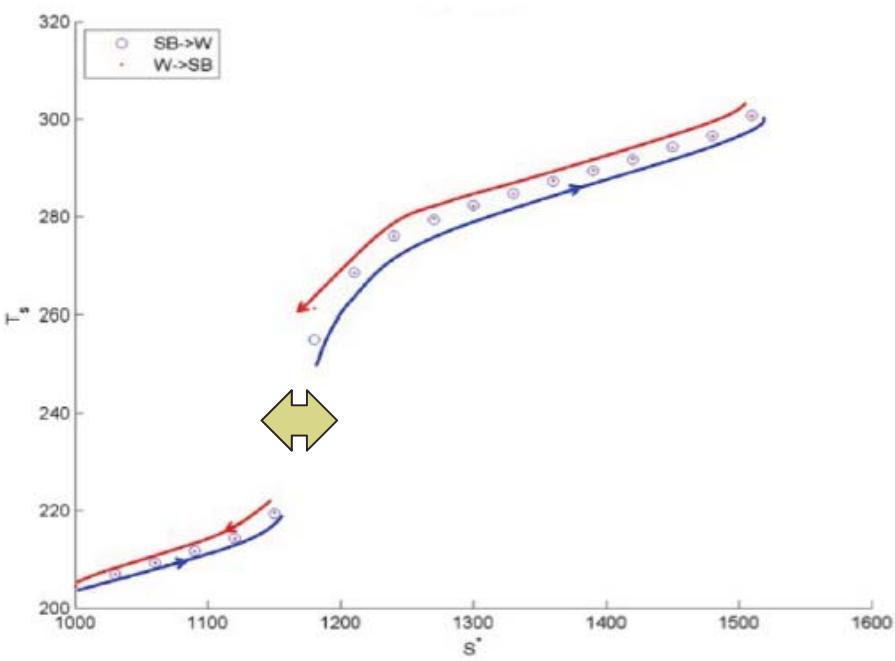




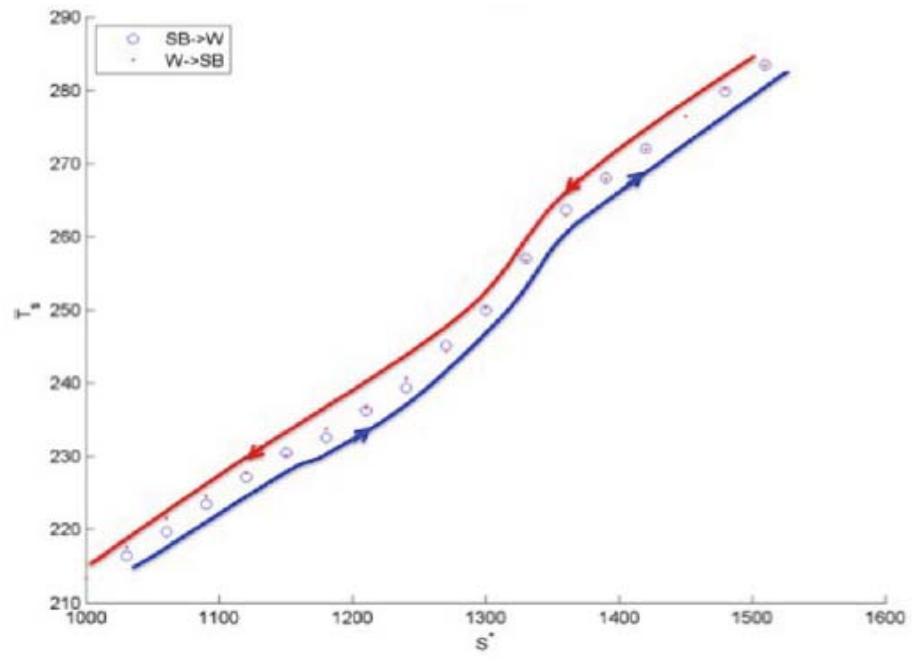
(a)



(b)

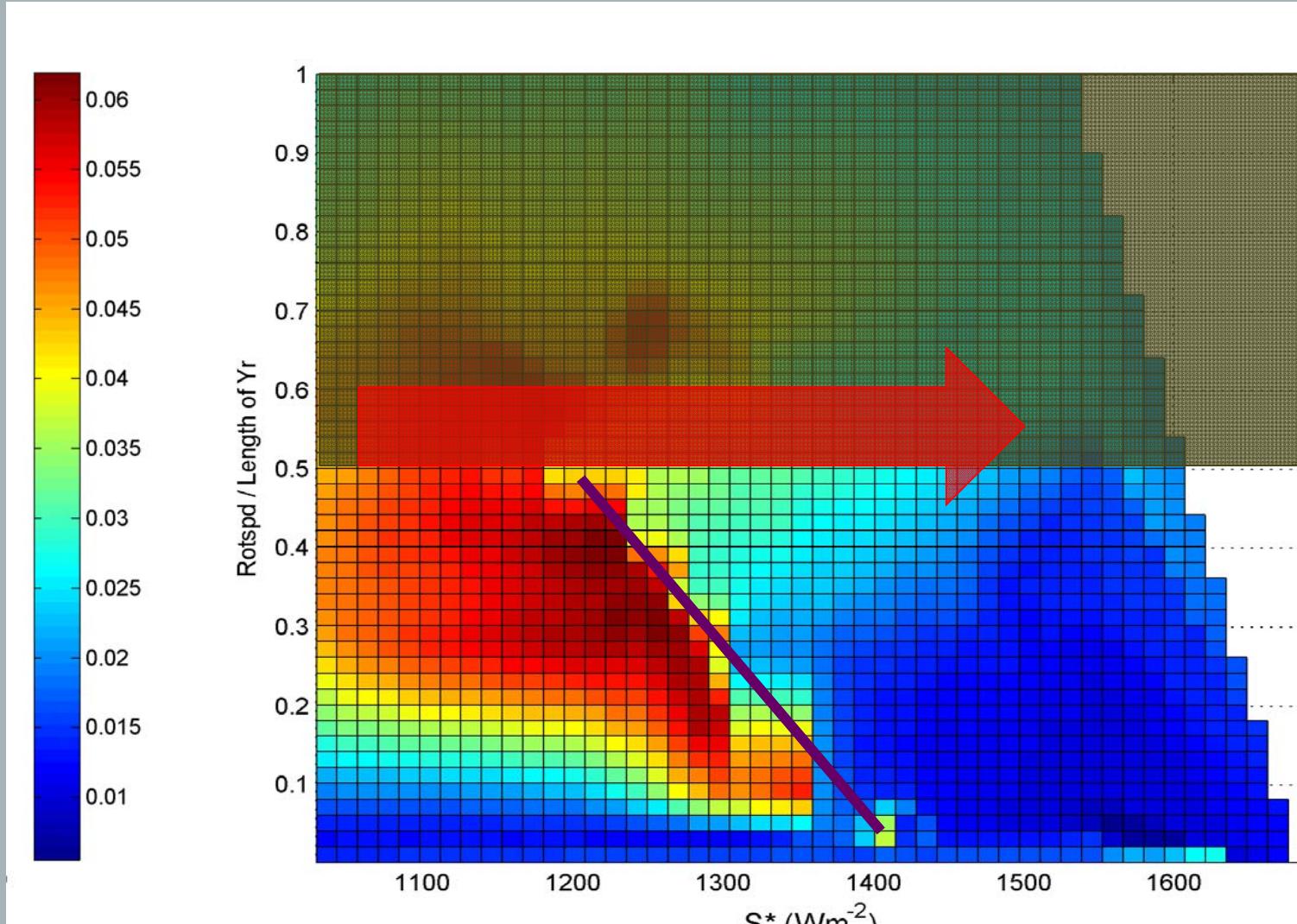


(c)

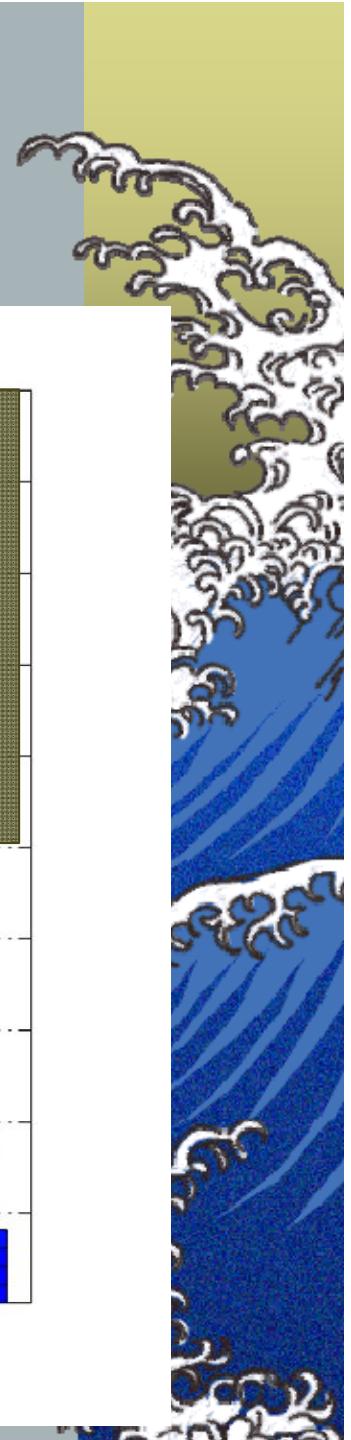


(d)

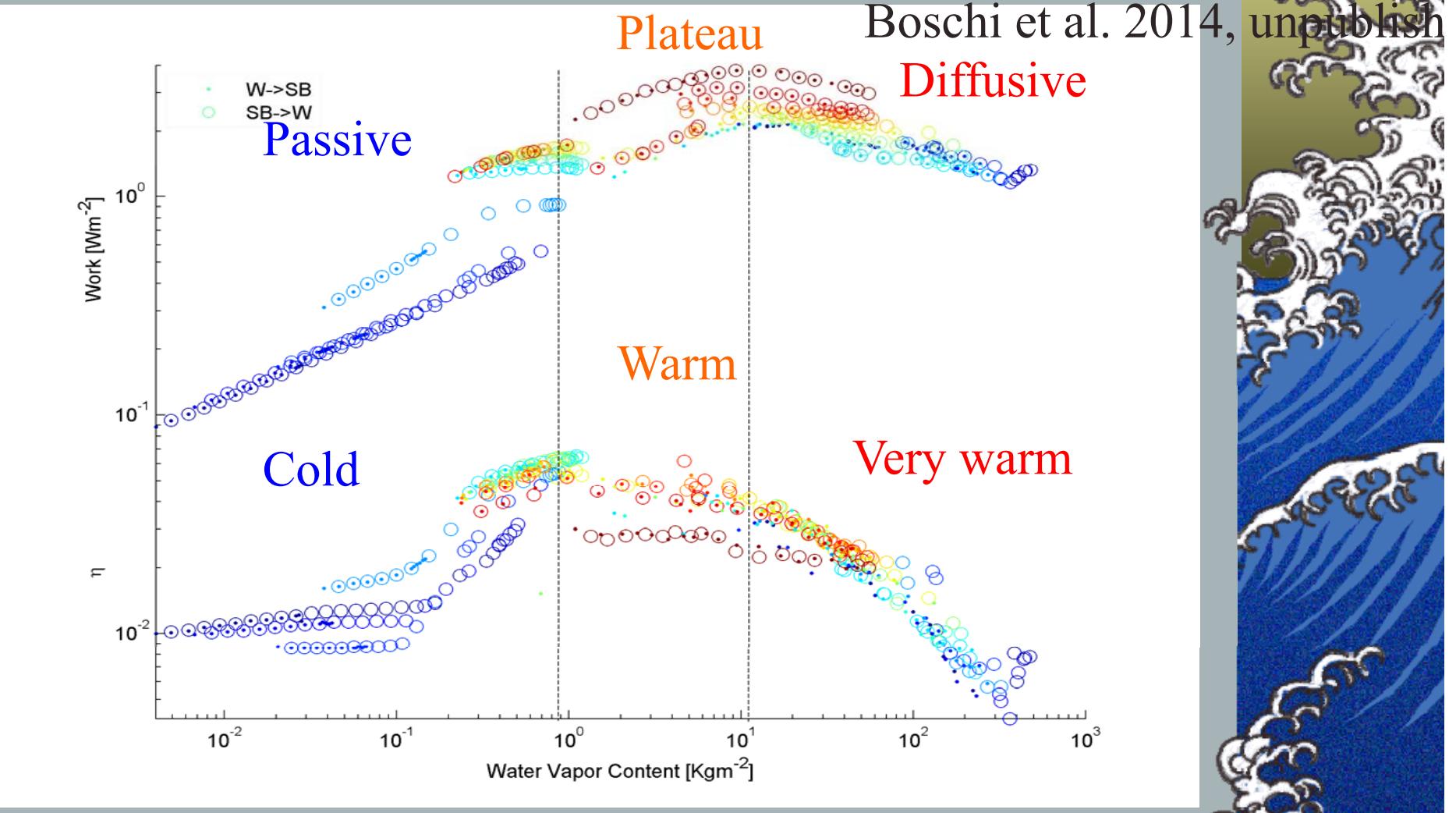
Bistability, Efficiency, Etc.



Boschi et al. 2014, unpublished



Water Vapour and Thermodynamics



Yr(Earth days)

360 180 36 12 4 3 2.4 2 1.8 1.6 1.5 1.333 1

Conclusions

- *Unifying picture connecting Energy cycle to EP;*
- *Simplified 2D formula for studying GCMs*
- *Snowball hysteresis experiment*
 - *Mechanisms involved in climate transitions;*
 - *Analysis of the impact of $[CO_2]$ increase*
 - *Generalized set of climate sensitivities*
 - *Analysis of impact of change on l.o.y.*
- ***Many challenges ahead:***
 - *Analysis of GCMs performance*
 - *Melancholia/Edge States*
 - *Multiscale, coarse graining effects*



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